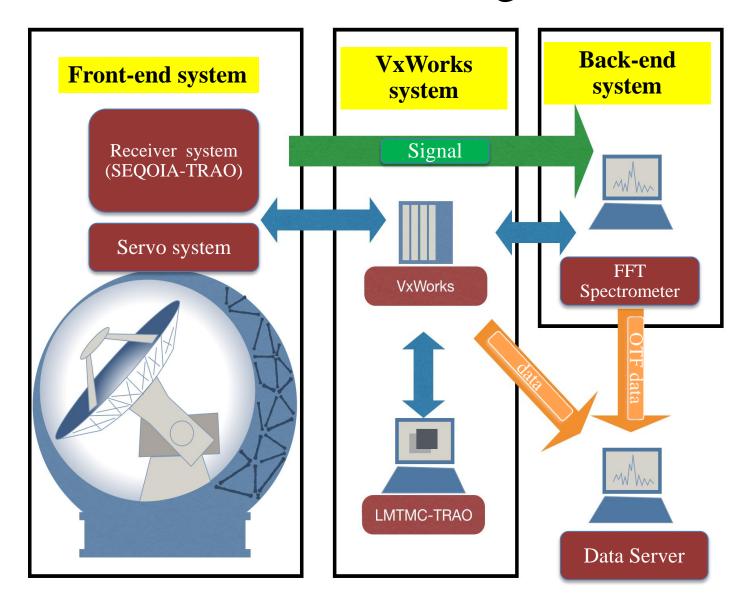
Upgraded TRAO and Its Performance

Chang Won Lee & TRAO team Korea Astronomy & Space Science Institute



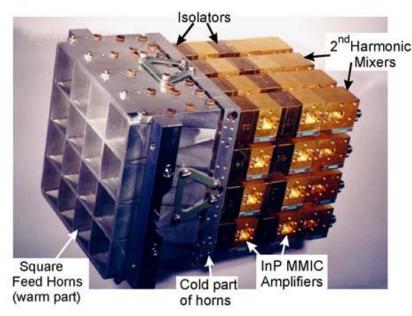
TRAO Block Diagram



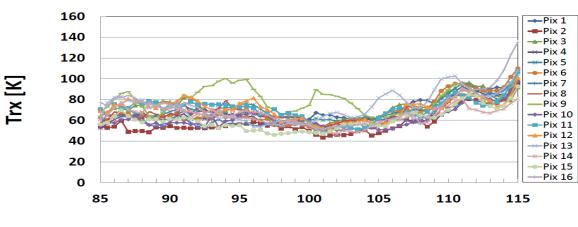
By H. Kang et al.

New receiver system

- •New Multi-beam receiver system (SEQUOIA-TRAO) imported from U. Mass, 16 pixel MMIC preamplifiers in a 4x4 array
- •Working frequency band : 85 115.6 GHz
- •Trx = 60-80 K for 86 110 GHz, 80-110 K for 115GHz



By C.H. Lee et al.



Frequency [GHz]

New Back-end systems

•New FFT spectrometer

- Full spectra bandwidth: 125 MHz (~375 km/s at 100 GHz)
- Best spectral resolution: ~15.5 kHz with 8192 channels (~0.04 km/s at 100 GHz)



→ With two 2nd LOs in a parallel mode set-up, *two molecular lines (maximum 15GHz apart) can be simultaneously observable*. A single observation will enable us to get the data of 16 beams in 2 lines simultaneously, with the spectral bandwidth of ~60MHz.

TRAO Observing mode: OTF

- On-the-Fly (OTF) mode (+ a simple position switching mode)
- Useful for mapping large area > 10' x 10'
- e.g.) Mapping 15' x 15' area which corresponds to a factor ~6 wider than 6' x 6' mapping area only requires twice the time for mapping area of 6' x 6', for the same sensitivity!

Mapping Area	Area Fraction (AF)	Time Fraction (TF)	Mapping Efficiency(AF/TF)
15' x 15'	6.25	2	3.1
12' x 12'	4	1.7	1.3
10' x 10'	2.8	1.4	2
6' x 6'	1	1	1

Performance

• (Best) System temperature

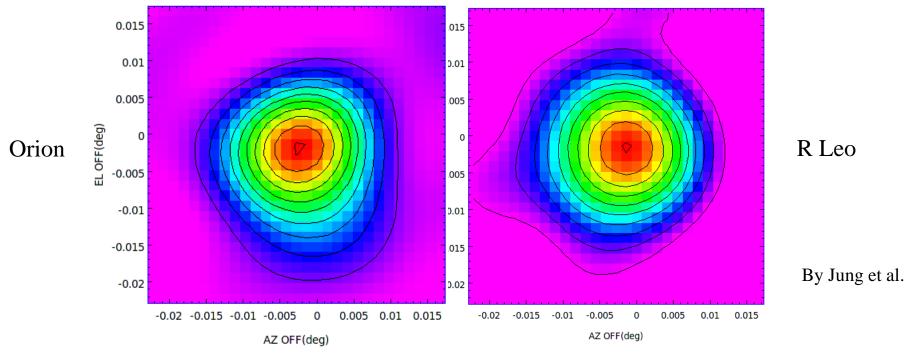
Mol. Line. (Freq. GHz)		N₂H⁺1-0 (93.176)	SO (99.2999)	NH ₂ D (110.154)	C ¹⁸ O 1-0 (109.782)		¹² CO 1-0 (115.271)
T _{sys} (K)	160	160	240	230	190	210	400

By Chung et al. and A. Soam et..al

•Pointing Accuracy ~10"

By Jung et al.

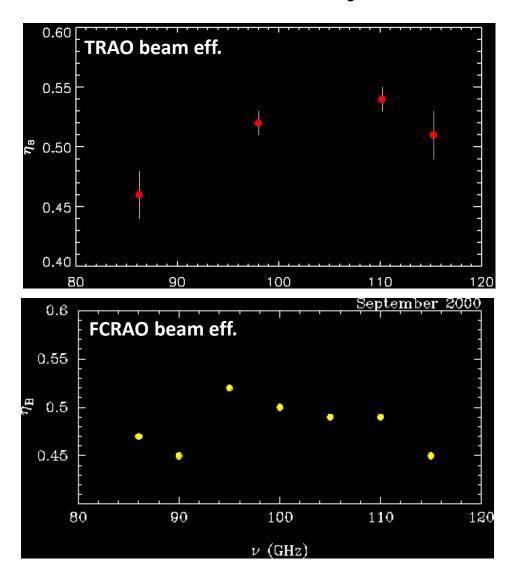
Beam Parameters



Frequency (GHz)	86.243	98.000	110.201	115.271
$\vartheta_{\scriptscriptstyle B}$ (")	60	53	47	45
η _Α (%)	39±2	44±1	46±1	43±2
η _B (%)	46±2	52±1	54±1	51±2

- Beam size from SiO Maser observations for Orion (left) and R Leo (Right) at 86.243 GHz.
- Beam efficiencies from Venus and Jupiter continuum observations

Beam Efficiency



By Jeongng et al.

TRAO Internal –Risk-Shared Observations (Jan. – April, 2016)

TRAO Key Science Programs

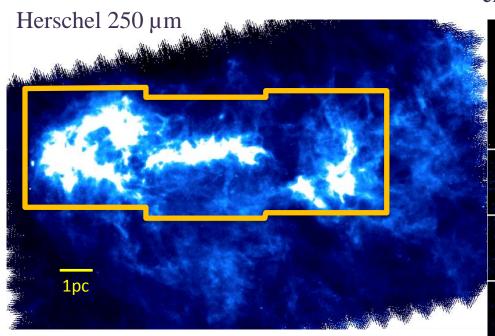
- 1. TRAO Multi-beam Legacy Survey of Nearby Filamentary Molecular Clouds (NFMC; CW Lee)
- 2. Mapping turbulent properties of star-forming molecular clouds down to the sonic scale (TPMC; JE Lee)
- 3. TRAO Observations of Planck cold clumps (TOP; T Liu)

General Programs

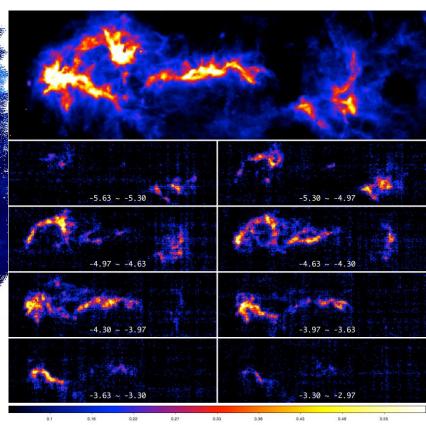
- 1.Molecular line diagnostics of the radiation-driven implosion mode of triggered star formation on birght-rimmed clouds (BRC; A Soam)
- 2. Molecular Line Observations of the Far-Infrared Bright Supernova Remnants (FIRSNR; I Jeong)
- 3. A Search for High-mass Star Forming Region Near Isolated Small H II Region (HMSFR; S Kang)
- 4. CO Survey of Inner Bar Region of the Milky Way (Inner Bar; Y Lee)
- 5. Methanol Outflow of the Extremely Young Protostar V380 Ori NE (Methanol Outflow; M Choi)
- 6.TRAO Outer Galactic Planer Survey (TOGS; Y Lee)

TRAO Results — Filamentary clouds

C¹8O 1-0 integrated intensity map and channel maps

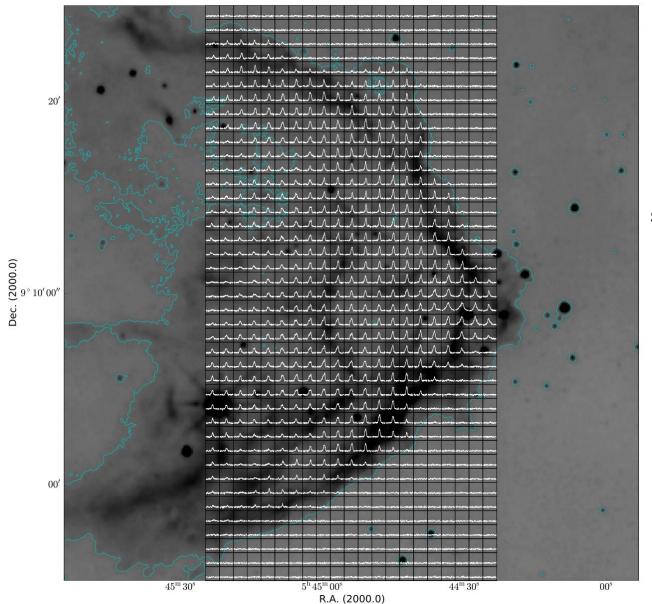


L1251 of Cepheus Molecular Cloud ~96'x32' in 0.155K rms (in T_A*),t_{int} ~20 hours, with simultaneous 13CO 1-0 observations



S. Kim et al. 2016 in prep.

TRAO Results - BRCs



SFO 18 CO(1-0) (+WISE $12\mu m$)

- Known to be in the no
-triggered star
formation category, but
now believed to be
"triggered" starformation region from
this observation.

A. Soam et al. 2016 in prep.

More for TRAO upgrade

- Radome will be replaced this October.
- New Servo system is planned to be replaced next year. Pointing stability and accuracy will be improved.
- Development of wideband (~2GHz) spectrometer is planned.

TRAO Open use

- TRAO Proposal call for general users for late Autumn 2016 and Spring of 2017 (till the end of April 2017)
- Dead line: September 30st, 2016 (by midnight KST)
- Maximum ~100 hours per proposal

Just visit TRAO web site and write your proposal for TRAO.

We welcome your proposal!