

The Virial Relation and Shape of Early-Type Galaxies

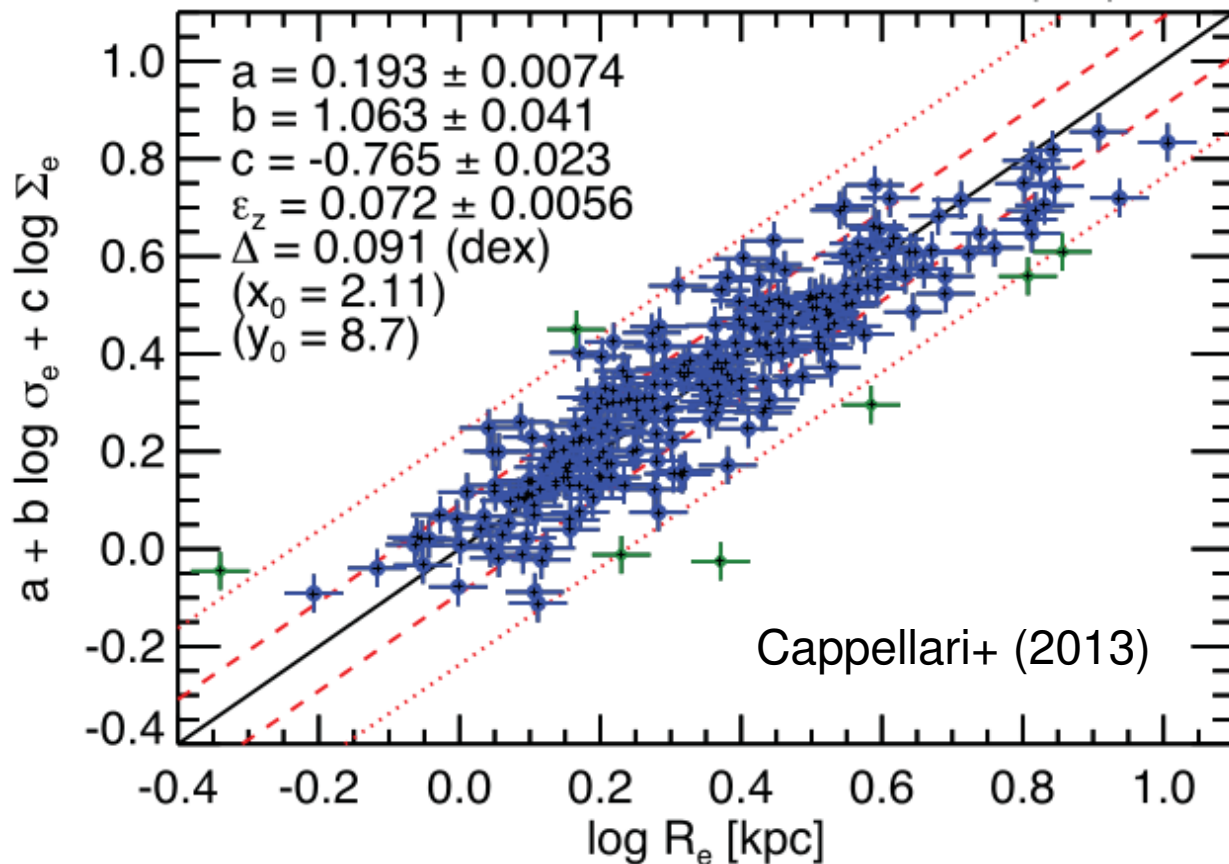
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The fundamental plane



$$\log R_e = a + b \log \sigma + c \log \Sigma_e$$

|
 eff. radius

/
 velocity dispersion

/
 surface brightness

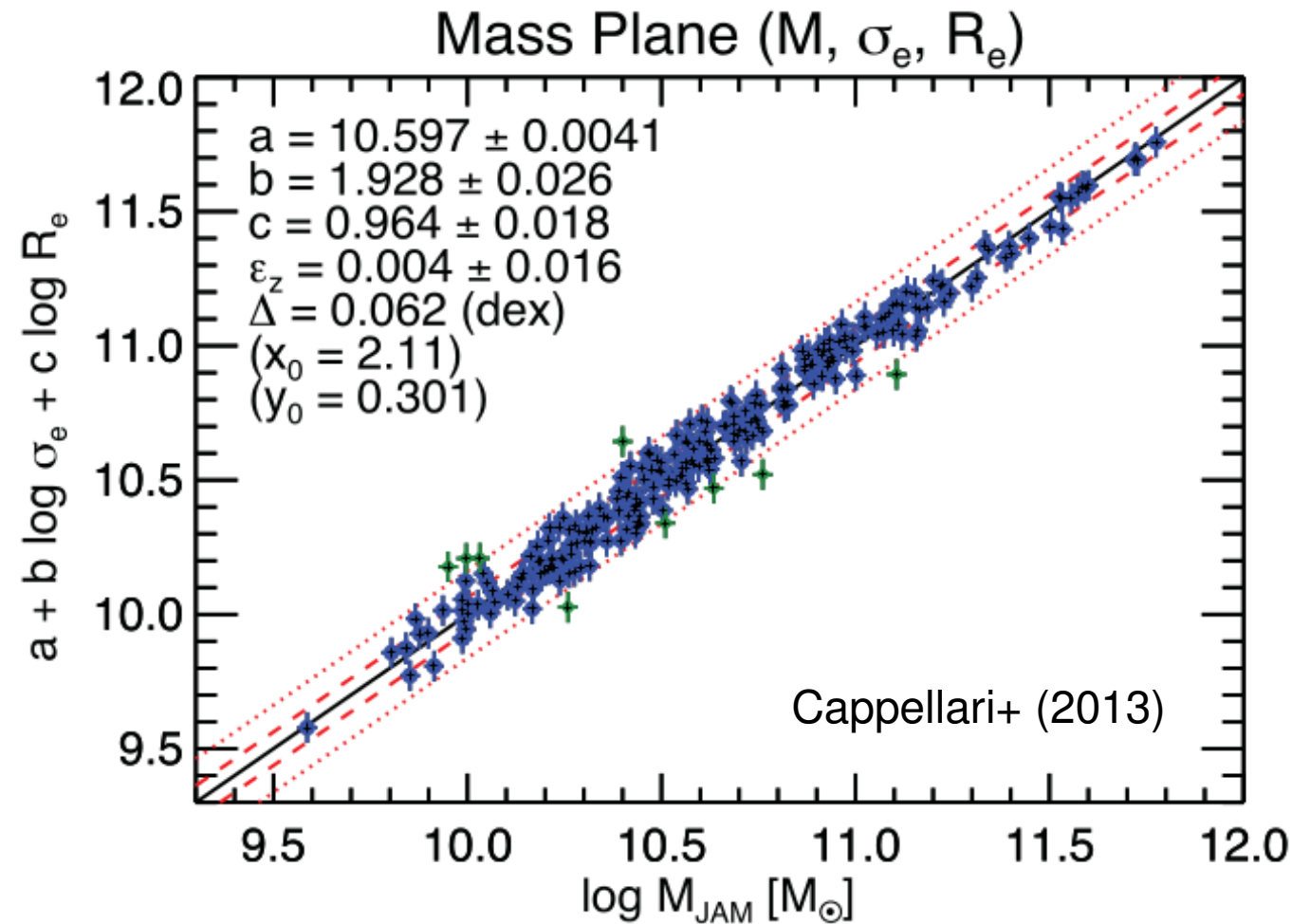
Expected: $b = 2$, $c = -1$

Observed: $b \approx 1$, $c \approx -0.8$

↳ Difference due to scaling of M/L with velocity dispersion

Dressler+ (1987); Djorgovski & Davis (1987); many others since then

The mass plane



Virial relation

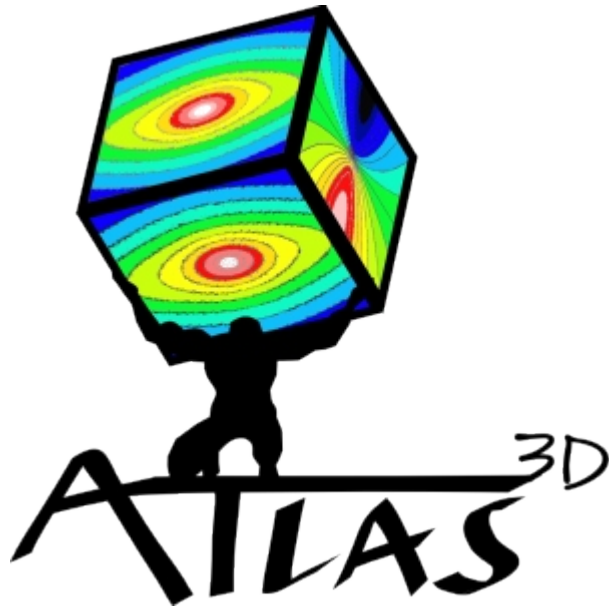
$$M = k_e \frac{\sigma_*^2 R_e}{G}$$

$$k_e \approx 4 \dots 5$$

Theory: $b = 2, c = 1$; observations agree? Or not?

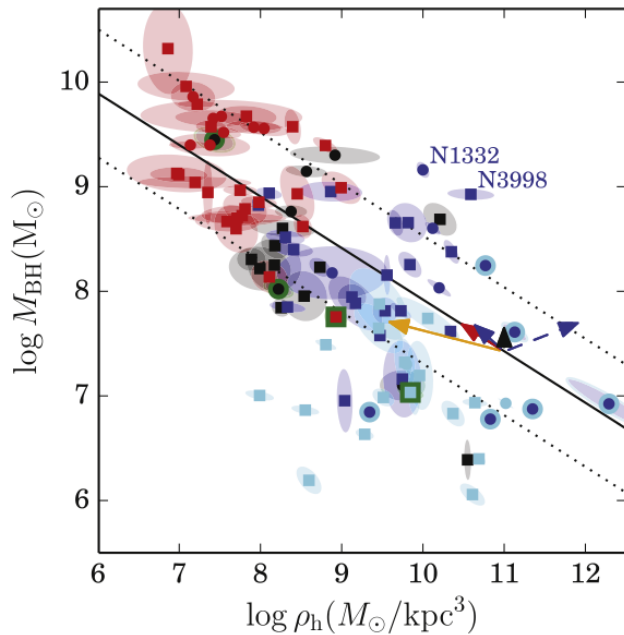
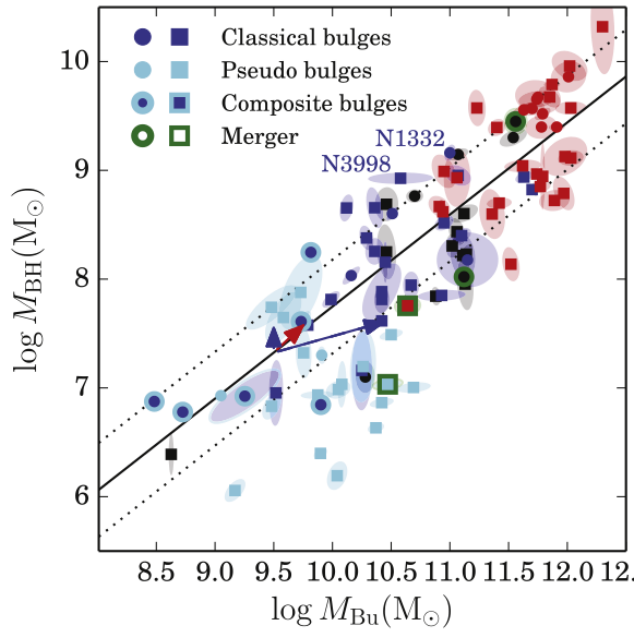
Bolton+ (2007); Hyde & Bernardi (2009); Cappellari+ (2013); Scott+ (2015)

The ATLAS^{3D} sample



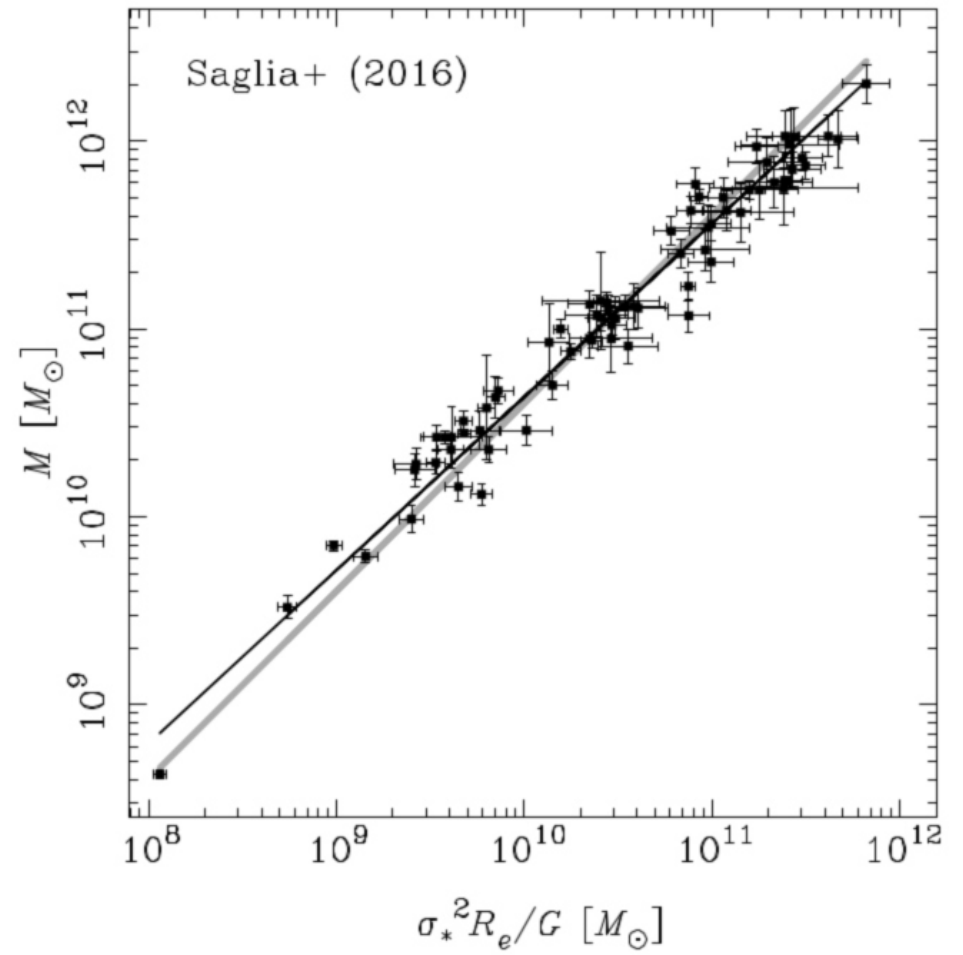
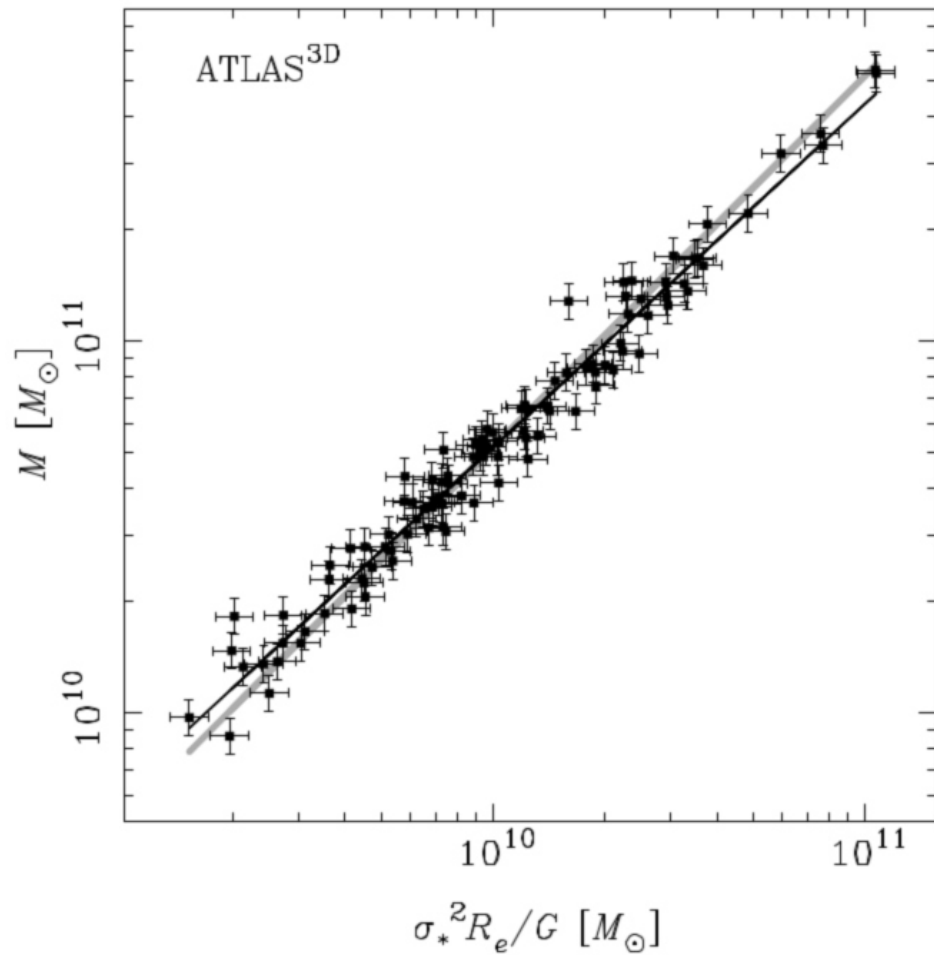
- ▶ 260 elliptical galaxies within 42 Mpc
- ▶ Volume-limited sample
- ▶ SAURON integral-field spectroscopy
- ▶ Masses and effective radii from modeling of surface brightness and velocity dispersion distributions in the sky plane
Jeans Anisotropic Multi-Gaussian Expansion (JAM) algorithm
- ▶ High-quality results for 101 galaxies
- ▶ Formal uncertainties:
 - 0.041 dex for effective radii
 - 0.021 dex for average velocity dispersions
 - 0.049 dex for galaxy masses

The Saglia+ (2016) sample



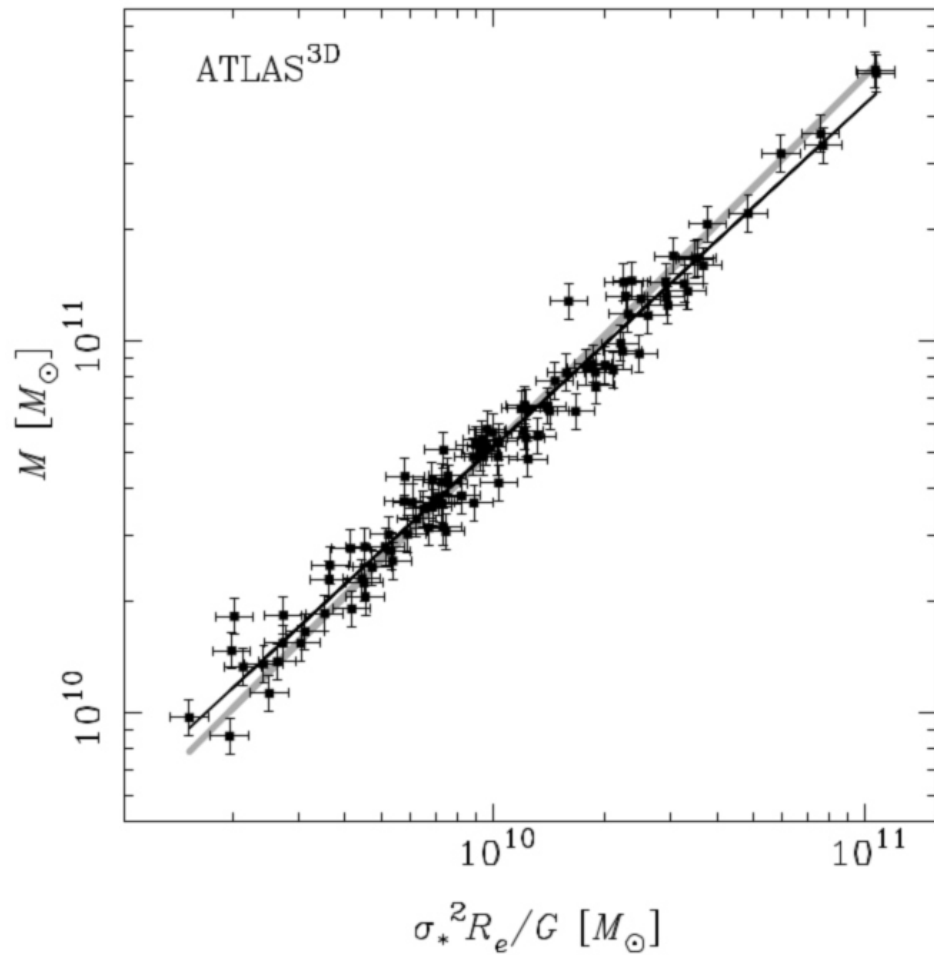
- ▶ 72 galaxies & bulges within 150 Mpc
- ▶ Aimed at black holes – galaxy relations, selected for large range in dispersion
- ▶ SINFONI integral-field spectroscopy, various archival data sets
- ▶ Masses and radii from photometry, various dynamical models
- ▶ Median formal uncertainties:
 - 0.096 dex for effective radii
 - 0.021 dex for average velocity dispersions
 - 0.083 dex for galaxy masses

The virial relation with R_e

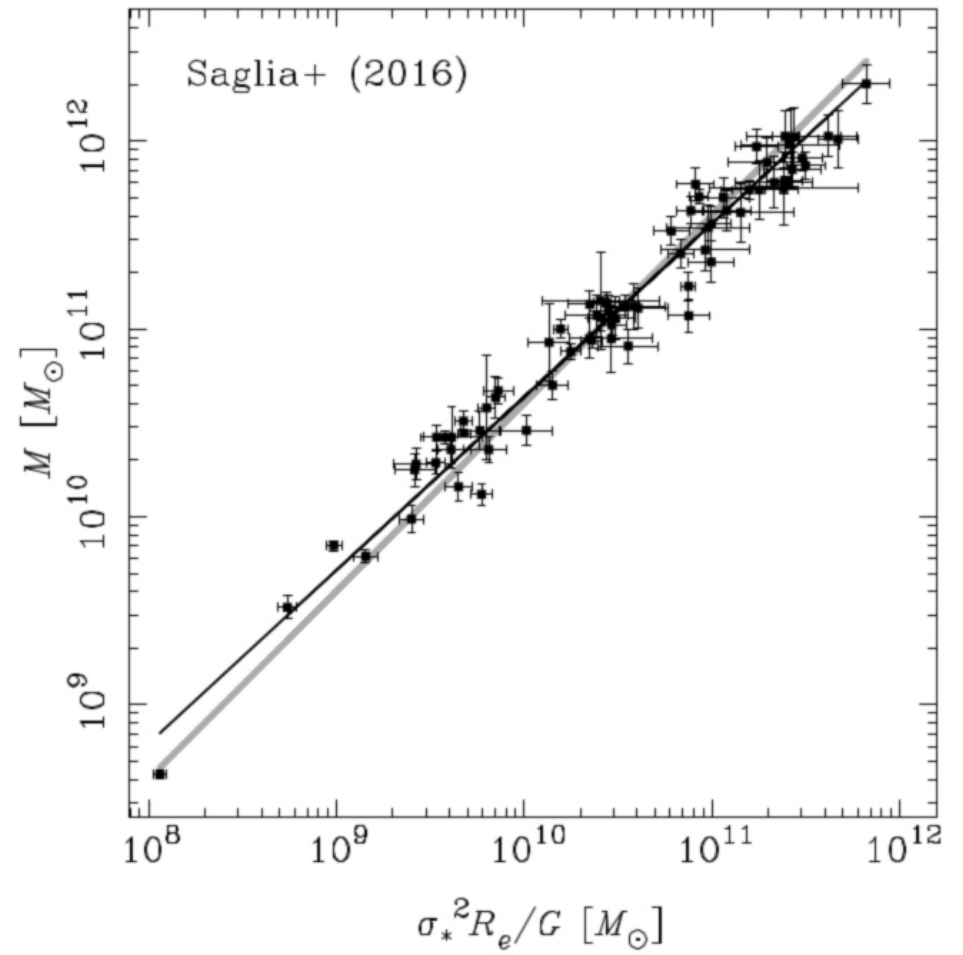


$$\log \left(\frac{M}{10^{11} M_\odot} \right) = x \log \left(\frac{\sigma_*^2 R_e / G}{10^{10.5} M_\odot} \right) + y$$

The virial relation with R_e



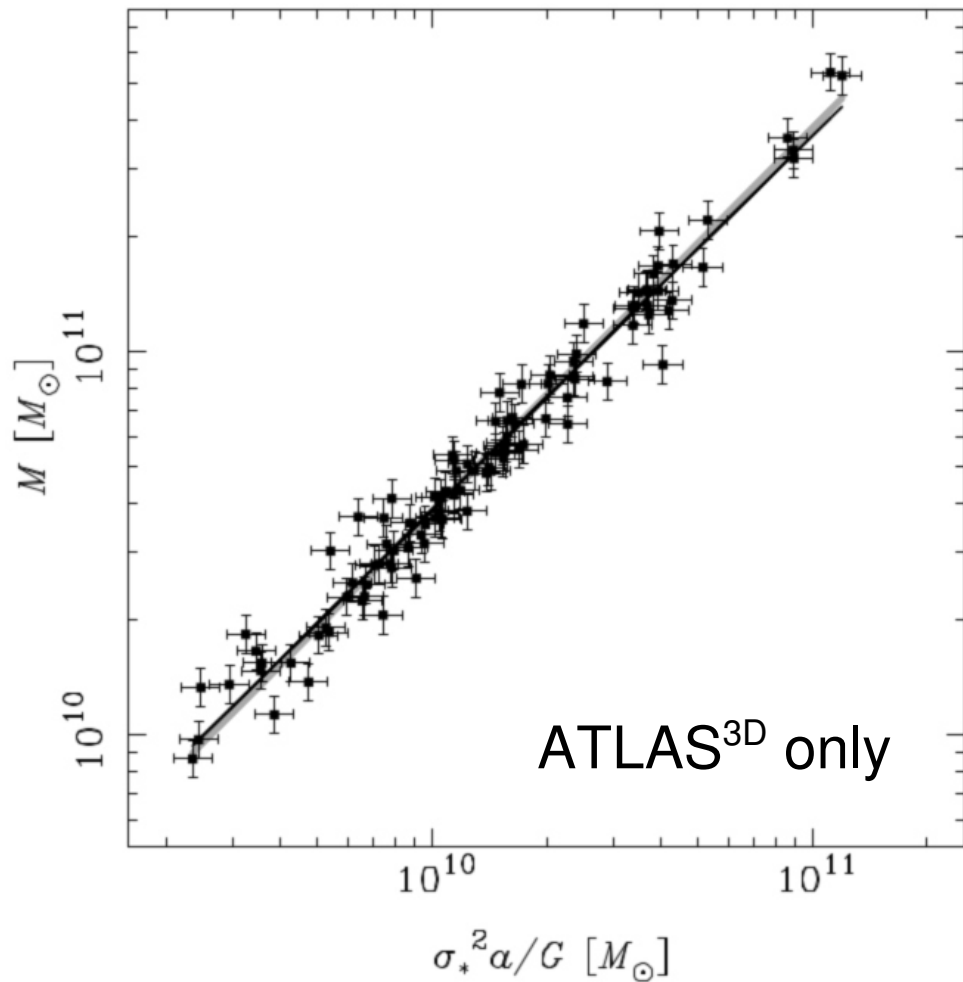
$$x = 0.924 \pm 0.016$$



$$x = 0.923 \pm 0.018$$

↳ Virial relation of ETGs is tilted

The virial relation with semimajor axis a



$$\log \left(\frac{M}{10^{11} M_\odot} \right) = x' \log \left(\frac{\sigma_*^2 a / G}{10^{10.5} M_\odot} \right) + y'$$

where the half-light ellipse obeys

$$R_e = \sqrt{ab}$$

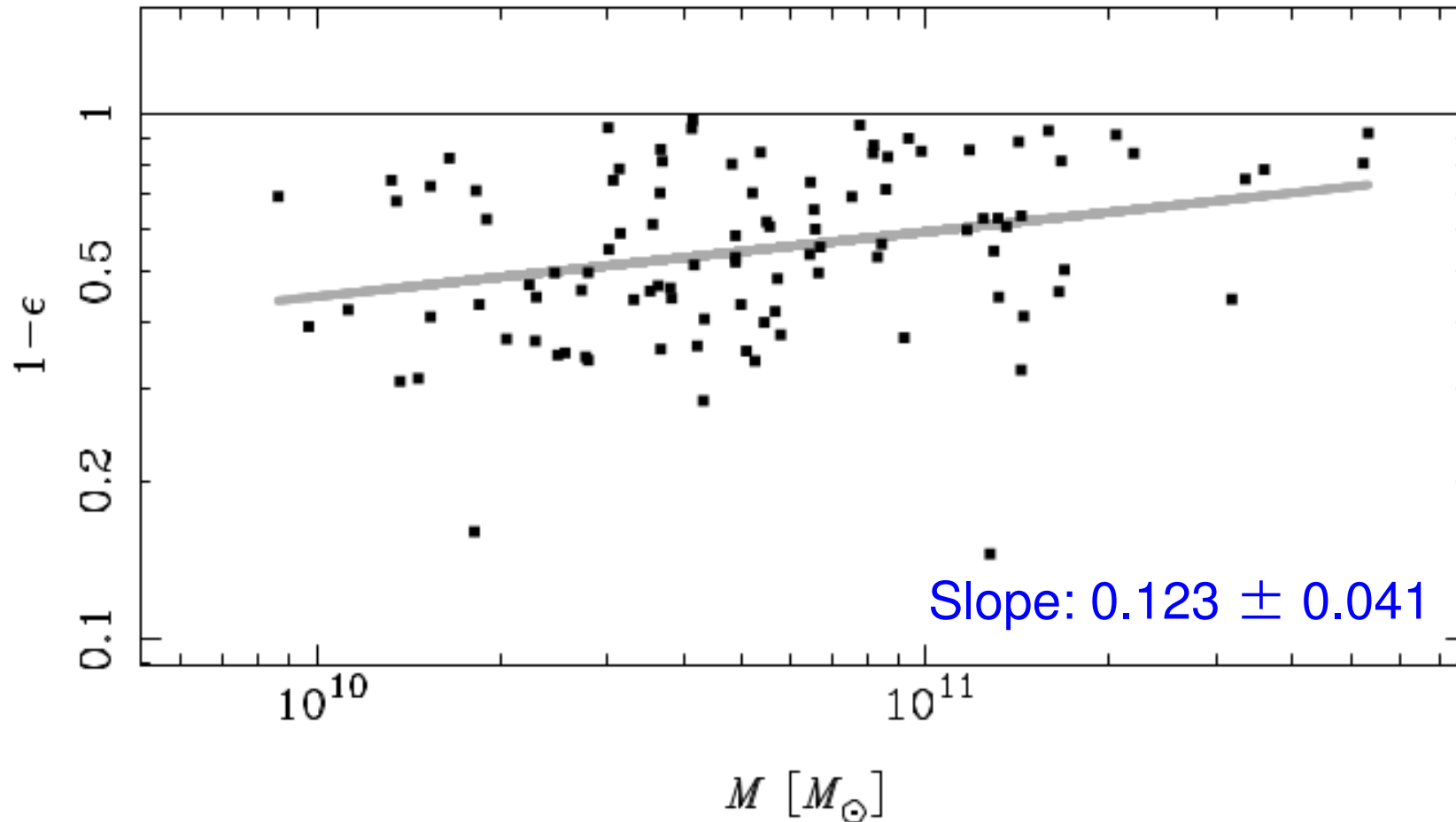
$$R_e = a\sqrt{1-\epsilon}$$

gives a new slope

$$x' = 0.976 \pm 0.018 \equiv 1$$

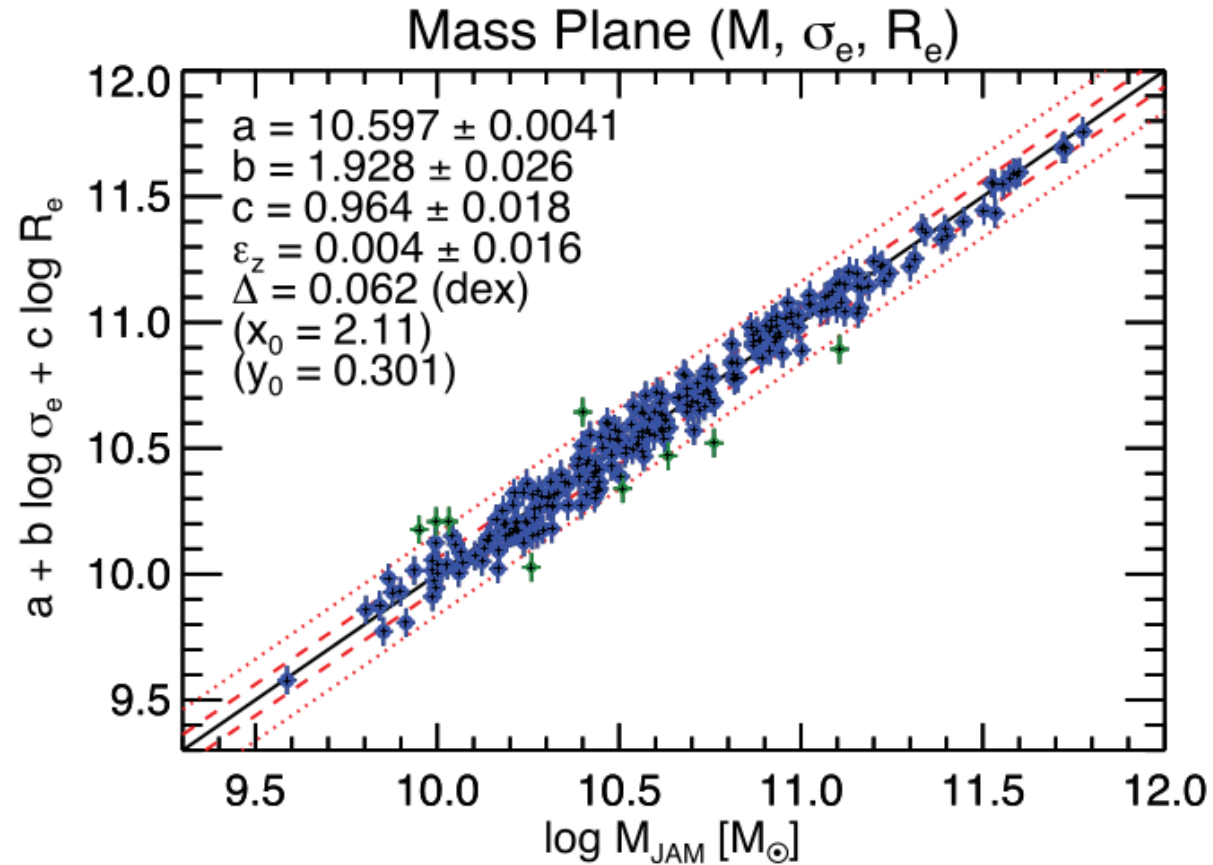
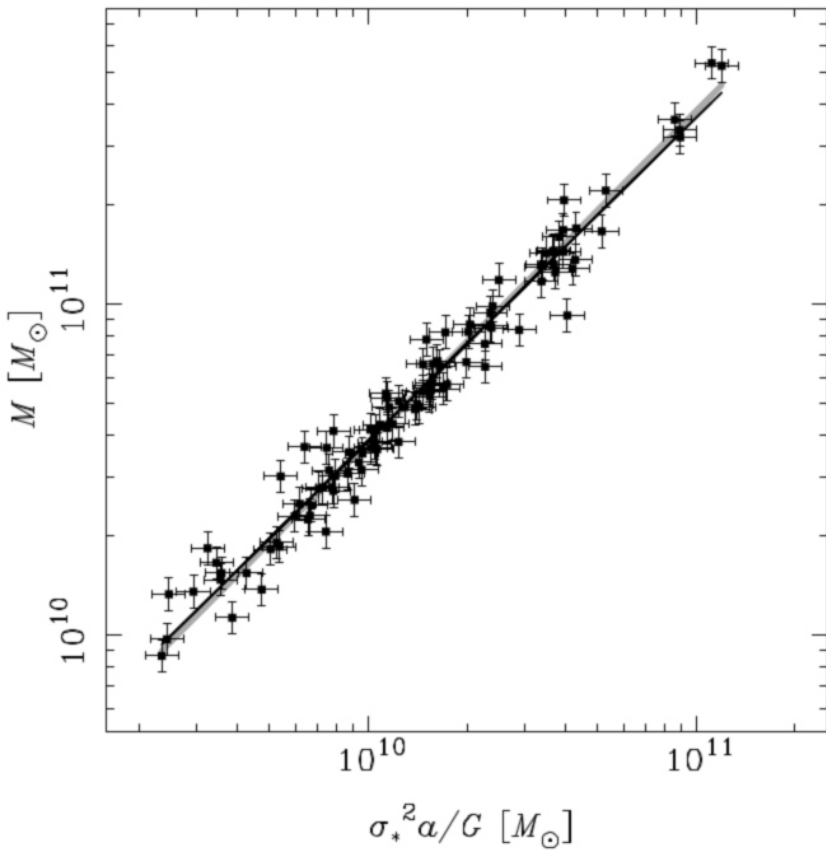
↳ Semimajor axis, not effective radius, is the true scale radius of ETGs

Ellipticity scales with mass



↳ Lack of very elliptical high-mass galaxies explains mass plane tilt

There is no mass “plane”

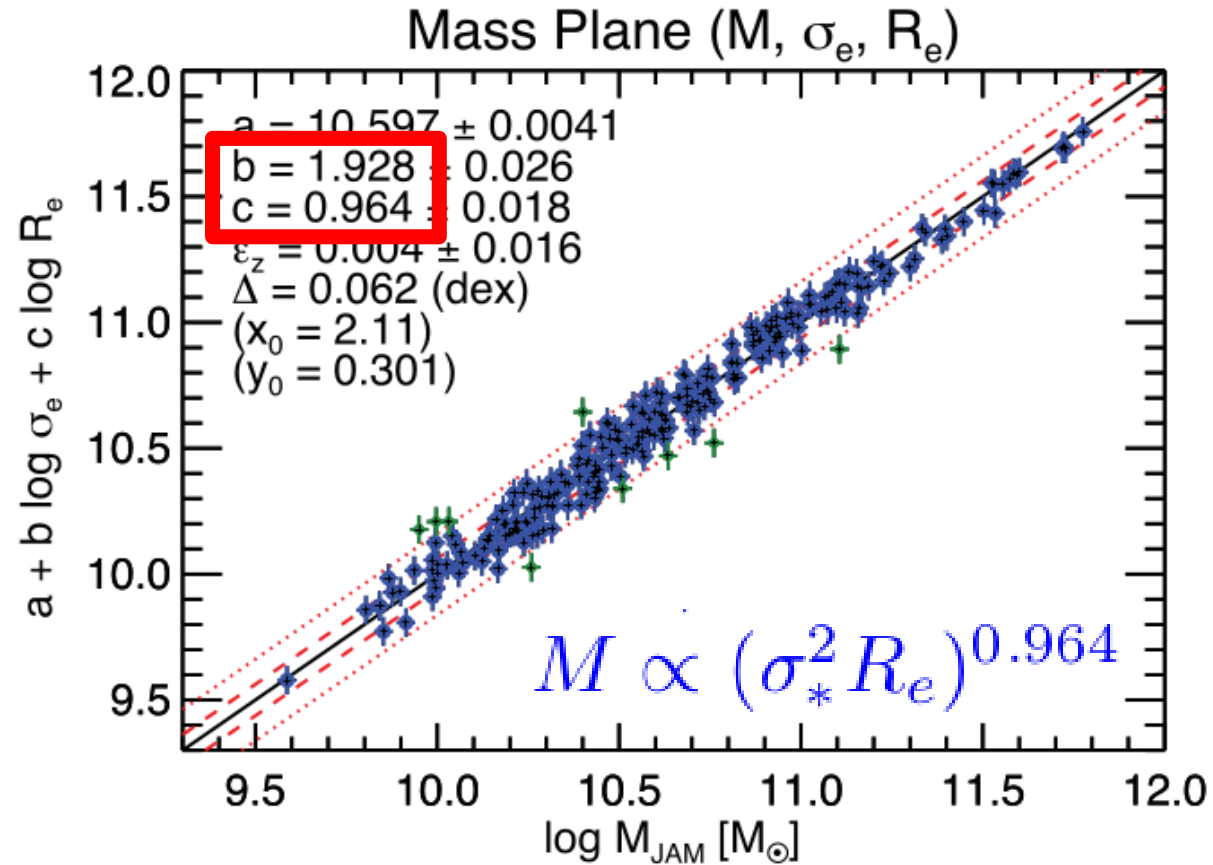
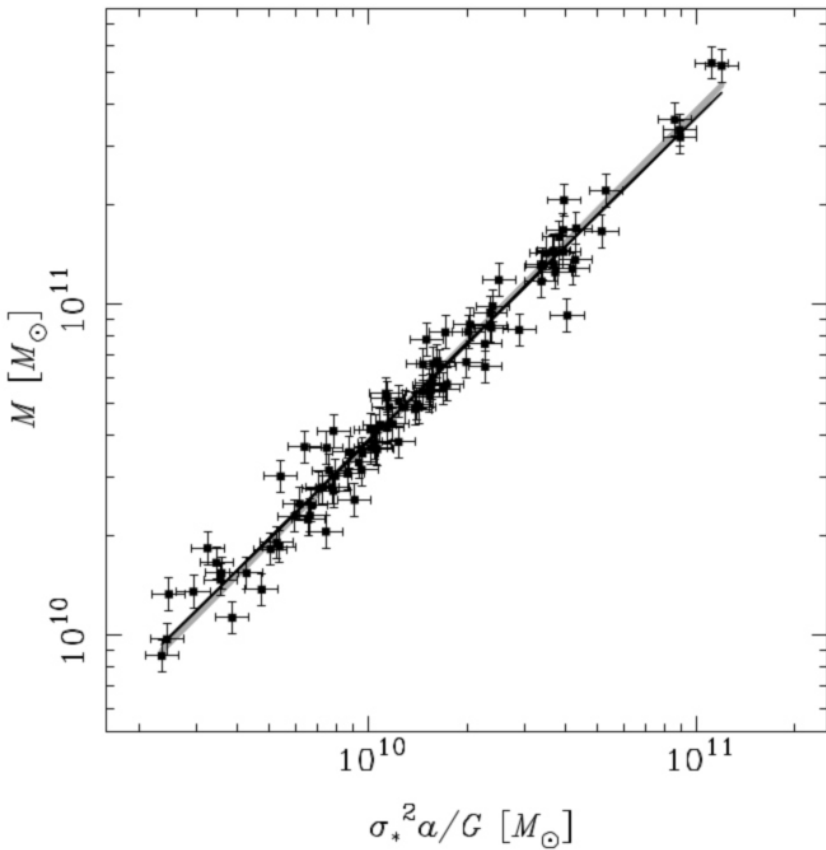


Best-fit virial relations agree with zero intrinsic scatter

↳ **2-parameter virial relation is sufficient (Occam's razor)**

↳ **3-parameter mass plane has too many parameters**

There is no mass “plane”

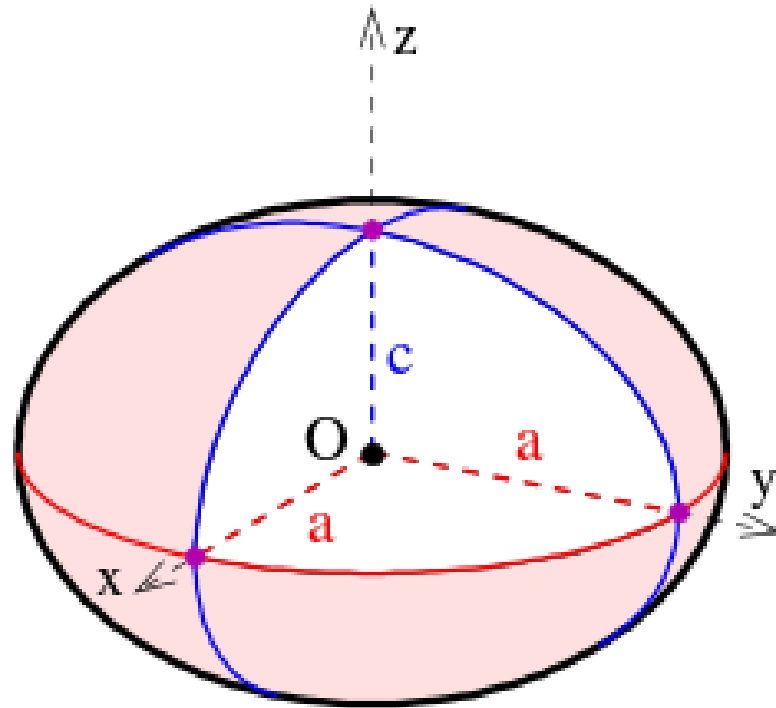


Best-fit virial relations agree with zero intrinsic scatter

↳ **2-parameter virial relation is sufficient (Occam's razor)**

↳ **3-parameter mass plane has too many parameters**

Local ETGs are axisymmetric and oblate



- ↳ If not, a were not a characteristic radius
- ↳ Agrees with dynamical modeling (Weijmans+ 2014)

Summary

- 72 & 101 ETGs from 2 recent surveys: ATLAS^{3D}, Saglia+ (2016)
- Using the virial relation with R_e shows a significant tilt
- Tilt disappears when using the semimajor axis a as scale radius
- Tilt is caused by a lack of very elliptical high-mass galaxies
- Virial relation fully determines ETG dynamics, no mass “plane”
- a is scale radius → local ETGs are (mostly) axisymmetric and oblate

Trippe (2016), JKAS, in press – arXiv:1609.07188

